

Project Details

ROSES ID: NNH14ZDA001N

Selection Year: 2014

Program Element: Physics of the Inner Heliosphere

Project Title:

Connecting theory and simulations of turbulence in the inner heliosphere with single-point measurements by Solar Probe Plus and Solar Orbiter

PI Name: Jean Perez

PI Email: atn@g.ucla.edu

Affiliation: University of New Hampshire

Project Member(s):

- Chandran, Benjamin D. G.; Co-I; University of New Hampshire

Summary:

The primary objective of this project is to develop new methods and tools to better understand the relation between single-point space-craft measurements and the three-dimensional physical properties of the solar wind in the inner heliosphere, which is critical for understanding and interpreting measurements from the upcoming Solar Probe Plus (SPP) and Solar Orbiter (SO) missions. The need for new methods derive in part because in this region, many of the assumptions that are usually made at higher heliocentric distances breakdown, thus requiring novel data-analysis methods in order to obtain meaningful quantitative predictions that can be compared to theoretical and numerical models of the solar wind, as well as other processes that impact life and society here on Earth. In order to achieve the goals of this project, we will produce synthetic data simulating measurements from instruments on board these missions by flying virtual probes in simulations of Alfvénic turbulence. The probe's trajectories will be chosen to simulate planned orbits of the Solar Probe Plus and Solar Orbit missions. We aim to develop new analytical methods to establish meaningful connections between the synthetic measurements and three-dimensional properties of the solar wind. The methods will be validated using the full three-dimensional structure of the underlying fields that are completely known and controlled in the simulations. The simulations will be performed inside a narrow magnetic flux tube that models a coronal hole extending from the solar surface into the solar corona to about 60 solar radii (0.3 AU), using a fully developed spectral element code called IRMHD (Inhomogeneous Reduced Magnetohydrodynamics). The IRMHD code is thus far the only code capable of simulating broad-spectrum non-compressive turbulence while resolving the inhomogeneity of the background profiles, including the outflow velocity, without approximating the nonlinear terms in the governing equations. The outcomes of this project are timely, given that they will be available when SPP and SO become operational, enabling future comparisons with SPP measurements that will contribute significantly to the science return of this pioneering mission.

Publication References:

Summary: no summary

Reference: Klein, Kristopher G.; Perez, Jean C.; Verscharen, Daniel; Mallet, Alfred; Chandran, Benjamin D. G.; (2015), A Modified Version of Taylor's Hypothesis for Solar Probe Plus Observations, The Astrophysical Journal Letters, Volume 801, Issue 1, article id. L18, 5 pp, doi: 10.1088/2041-8205/801/1/L18

Summary: no summary

Reference: Chandran, Benjamin D. G.; Perez, Jean C.; Verscharen, Daniel; Klein, Kristopher G.; Mallet, Alfred; (2015), On the Conservation of Cross Helicity and Wave Action in Solar-wind Models with Non-WKB Alfvén Wave Reflection, The Astrophysical Journal, Volume 811, Issue 1, article id. 50, 7 pp, doi: 10.1088/0004-637X/811/1/50